ysics" ~ 2005	del at HERA	APNIA / Spp	Remember : H1, Z. Phys. C74 (1997) 191 ZEUS, Z. Phys. C74 (1997) 207 Large interest in the community. 300 citations each, show that: 300 citations each, show that: Yes, BSM surprises might come from HERA !
"New Trends in HERA Ph Ringberg, 2 nd -7 th Octobel	the Standard Mo	Emmanuelle PEREZ SY & CEA-Saclay, DSM / D.	$\left(\frac{10^{-1}}{10^{-1}} + \frac{10^{-1}}{10^{-1$
	Beyond	DES	E. Perez

Searches for new currents affecting the DIS process - Charged current DIS - Neutral current DIS	 Aodel dependent searches for new particles : Since HERA is not an annihilation machine, the cross-section is small for (heavy) pair-produced new particles. The limits depend on the coupling of the new particle to SM ones, i.e. no "absolute" mass limits in contrast to LEP, Tevatron. In case of a discovery: information not only on M but also on this coupling Examples : Leptoquarks, Rp-violating SUSY, excited fermions. 	<mark>lodel independent searches for new physics :</mark> - Study SM processes with a low cross-section - investigate all possible final states, compare data to SM expectatio
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Introduction & Outline

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	d W bosons ith no WR atron	9 <mark>8</mark> : 788 GeV	07/10/05
new W' bosons	e^{\pm} , SM) ~ (1 \pm P_e) ES with polarised e a priori sensitive to right-handed vided that v _R is light and stable) polations to P _e = ± 1 consistent w seponding bound on M(V _R) if g_L = much below that reached at Tev much below that reached at Tev	vents per 5 GeV/c ⁶	$\int_{10^{-2}}^{10^{-1}} \frac{-W \rightarrow e}{-Wuttijet}$ 10 ⁻² - Other backgrounds 10 ⁻² 10 ⁻² 10 ² 10 ³ 10 ³ 10 ³ M _T (GeV/c ²)
CC DIS and r	CC D CC D Correction CC D CC D	Future HERA sensitivity <mark>might reach</mark> ~ 350 GeV (high lumi, high precision on polarisation measurement)	i.e. unlikely to discover a W _R at HERA-II - but nice textbook plot ! ^{E. Perez}

	con ! a dif cm (f_e=1)	¹⁶ cm (f ₆ =f ₄)	
2	o ~ 40000 GeV icision : ~ 3% at it 10000 GeV ² by statistical er lot of informat le : an and an at $an an an an at le : an an an an at $	2 - R= 0.7 * 10 - H 10 10 10 - H	
NC DIS at high Q	HERA Neutral Current 10 10 10 10 10 10 10 10 10 10	$\int_{10^{-7}}^{10^{-7}} \int_{10^{-7}}^{10^{-7}} \int_{10^{-7}}^{10^{-7}} \int_{10^{-7}}^{10^{-7}} \int_{10^{-7}}^{10^{-7}} \int_{10^{-7}}^{10^{-4}} \int_{10^{-7}}^{10^{-4}$	Perez 4 4
	qa/dQ ² (pb/6eV ²)		ய்



L and B together: Leptoquarks	ent symmetry between the lepton & quark sectors ? cancellation of QED triangular anomaly ?	appear in many extensions of SM pp ep ep ep ect lepton & quark sectors	rr or Vector color triplet bosons \overline{q} \overline{q} \overline{la} \overline{a}	ssified by Buchmuller, Ruckl and Wyler according to their quantum nbs tumptions : LQs couple only to SM fermions and bosons pure chiral couplings family diagonal couplings BRW model : if LQs \subset isospin multiplet are degenerate : $\begin{cases} B(LQ \rightarrow lq) &= 1 \text{ or } 1/2\\ B(LQ \rightarrow \nu q) &= 0 \text{ or } 1/2 \end{cases}$	me of) BRW assumptions relaxed \Rightarrow "generic models" $\Rightarrow \beta_{1} = \beta$ (LQ $\Rightarrow 1 q$) and $\beta_{v} = \beta$ (LQ $\Rightarrow v q$) free parameters	
	Apparent <mark>sym</mark> ı Exact cancella	 LQs appear i (enlarged gauges) Connect lept 	 Scalar or Ve Carry both L 	 Classified I Assumption BRW mo 	\rightarrow (Some of) $\rightarrow \beta_1 =$	c L

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07/10/05

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<mark>In case a s</mark> angular dis	tributions		oinatoria	l backgro	und, large	y looking (e acceptai	u ne
Further de F = 0 or Chiral co Couples 1	eterminati 2 ? uplings ? to v ?	on of son Compar Play wi Easy t	ne of the re rates i th polaris o see sinc	LQ quan n e ⁻ p anc sation of ce good S	tum numb 1 etp lepton be 5/B in vj o	sers : cam channel	
	$S_{0,L}$	$S_{1,L}$	$ ilde{S}_{0,R}$	$S_{0,R}$	$S_{1/2,L}$	$ ilde{S}_{1/2,L}$	$S_{1/2}$
$S_{0,L}$		β_{ν}	P_e	P_e			
$S_{1,L}$	$eta_{ u}$		P_e	P_e			
$ ilde{S}_{0,R}$	P_e	P_e		P_p		a / a	
$S_{0,R}$	P_e	P_e	P_p				
$S_{1/2,L}$						P_p	P_e
${ ilde S}_{1/2,L}$		e+	$/e^{-}$		P_p		P_e
$S_{1/2,R}$					P_e	P_e	

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	2	k prospect	S
Taking into acc Can we still hop	count our bounds (ob [.] pe to observe a LQ s	tained from 100 ignal at HERA-I	pb ⁻¹ of e ⁺ p data) : [?
e.g. for a F=0 L	Q which would manit	fest itself as a r	esonance i.e. M < ~ 280 GeV:
Total e ⁺ p luminosit	ty $\sigma_{prod} = current limit$	σ _{prod} = 0.5 * limit	(mu childe estimates denend
350 pb ⁻¹	~40	~ 2.5 c	a bit on the LQ mass, numbers
700 pb ⁻¹	* 5 ດ	~ 3.5 a	to be taken with care)
(i.e. H1 + ZEUS)			
ZEUS	5 Prelim, 05 e ⁻ p, 45 pb ⁻	yes if	the coupling is not
	- -	too fα	r from our present limits.
D\dq)	2 7 7 7		s might be significant only by
 	, a a a a a a a a a a a a a a a a a a a		starting in this direction
	IC (prel.) 05 e ⁻ p (45.1pb ⁻¹)	(comt	oining HERA-I limits)
10 ⁻⁵ SM (ZEL 10 ⁻⁶	US-S) P=-25.9%	NB : so fo	r HERA-II still looks very
2	10 ³ 10 ⁴ Q ² (c	àeV ²) much SM-	like
E. Perez	-	10	07/10/05

New physics in eq & DGLAP ? at has been done so far : - the SM expectation is taken from global fits, which include high Q^2 HERA data (94 data for CTEQ5) and other eN dat - checked that the SM exp. does not change much when using older fits (but the error on it does change a bit) kely that a significant NP effect might have "faked" DGLAP in fi real datasets (eN, μ N, ν N) uld be nice to quantify still: a combined fit of pdf's + NP terms. uld be nice to quantify still: a combined fit of pdf's + NP terms. the recent H1 fit of pdf's + EW parameter $\rightarrow J$. Meyer's talk for (high mass) LQ exchange: no full NLO calculation for DIS + already a LO fit would be informative.	
 What h What h Unlikely Unlikely Would Would Cf the Cf the but alre 	

<u>11</u> <u>10</u> are competitive with those from rare	501070 Gev 1 3 L 1 B 1 C C C C C C C C C C C C C C C C C	ng LQS erent generations. ¹⁰ ¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻²⁴⁰ ⁻²⁴⁰ ¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻²⁴⁰ ⁻²⁴⁰ ¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻²⁴⁰ ¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰ ⁻¹⁰	Fermions of diff from $\mu\mu$ and ely suppressed ely suppressed ction. $\mu(\tau)$ for $M < \sqrt{s}$ $\mu(\tau)$ for $M < \sqrt{s}$ eq $\lambda_{lq} / M^2 LQ$ $B_d \rightarrow e\mu$ (Be $B_d \rightarrow e\mu$ (Be are compet	In Fla pling to f s. kgrounds kgrounds bIS large inal select inal select inal select inal select $\mathbb{A}e, \lambda$ \mathbb{A} $\mathbb{A}e, \lambda$ $\mathbb{A}_{\mathbb{Q}}^{2}$ \mathbb{A} types \mathbb{Q} types \mathbb{Q}^{1} \mathbb{Q}^{1} \mathbb{Q}^{1} \mathbb{Q}^{1} \mathbb{Q}^{1}	Cpto Cs councesse rocesse CC Dac $CC Dac CC DaccC Dac cC DaccC Dac cC DaccC $	b LFV p LFV p $\mu_{\tau} \tau (v)$ $\mu_{\tau} (v)$ $\mu_{\tau} (v)$ $\mu_{\tau} (v)$ $\mu_{\tau} (v)$	Para and the state of the state	Possibilit Could lec Could lec q_1 q_1 For very in units in units ed $\rightarrow \mu b$
	Instraints	mples where DIS co	Several example	$B \rightarrow \tau \bar{e} X$ 14	$B \rightarrow \tau \bar{e} X$ 14	*	3 2	eb → μs
eb $\rightarrow \mu s$ 32 * $B \rightarrow \tau \bar{e} X$ $B \rightarrow \tau \bar{e} X$ 14 14 14 $\rightarrow \mu s$ Several examples where DIS constraints	11 e	d A		1.9	1.8	=		•
$eb \rightarrow \mu s \begin{bmatrix} 32 \\ 32 \\ 14 \end{bmatrix} * \begin{bmatrix} 1.8 \\ 1.9 \\ 14 \\ 14 \end{bmatrix} \xrightarrow{1.9} bd \overrightarrow{-} e\mu \text{ (pelle)} d \lambda_{11} e \\ \hline bd \rightarrow r\bar{e}X \\ \hline d \lambda_{11} e \\ \hline d \lambda_{11} h \\ \hline d \lambda_{11} \lambda_{12} \lambda_{12}$		trained by LQ	Better cons	$\begin{array}{c} B \rightarrow \mu \bar{e} \\ 0.4 \end{array}$	$\begin{array}{c} B \rightarrow \mu \bar{e} \\ 0.4 \end{array}$	*	13	ed → ub
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	ц Ц	Р У2		2 types	for ≠ L(TeV ⁻² ,	of -	in units
in units of TeV ⁻² , for \neq LQ types ed $\rightarrow \mu b$ $13 * B \rightarrow \mu \epsilon B \rightarrow \mu \epsilon B \rightarrow \mu \epsilon B B$	0.7 0.8 0.9 1 D	0 0.1 0.2 0.3 0.4 0.5 0.6 (, / MLQ ²	eq1 _Å µq2	ds on λ	uno	ZEUS b
ZEUS bounds on $\lambda_{eq1} \lambda_{\mu q_2} / M_{LQ}^2$ in units of TeV ⁻² , for $\neq LQ$ types ed $\rightarrow \mu b$ $13 \times B^{-\mu \overline{e}}$ $B^{-\mu \overline{e}^{-\mu \overline{e}}$ $B^{-\mu \overline{e}}$ $B^{$	501070	ZEUS, hep-ex/0	eg ^J Ig / M ² LQ	v no spur	es : por	ge mass	larg	For very
For very large masses : bounds on $\lambda_{eq} \lambda_{lq} / M^2 L_Q$ ZEUS bounds on $\lambda_{eq1} \lambda_{\mu q2} / M_{LQ}^2$ in units of TeV ⁻² , for $\neq LQ$ types ed $\rightarrow \mu b$ $13 \times \frac{B \rightarrow \pi \bar{e}}{1.8} \frac{B \rightarrow \pi \bar{e}}{1.9}$ Better constrained by $\frac{\overline{D} - \lambda_{23} - \mu}{d - \lambda_{11} - e}$ eb $\rightarrow \mu s$ $32 \times \frac{B \rightarrow \pi \bar{e}X}{3} \frac{B \rightarrow \pi \bar{e}X}{3} \frac{B \rightarrow \pi \bar{e}X}{14} \rightarrow \text{Several examples where DIS constraints}$	- T - T		$\mu(\tau)$) for M < \sqrt{s}	LQ, Je, J	N) uo pa	obtaine	nts	Constrai
Constraints obtained on $(M_{LQ}, \lambda_{e}, \lambda_{\mu}(\tau))$ for $M < \sqrt{s}$ For very large masses : bounds on $\lambda_{eq} \lambda_{lq} / M^2 L_Q$ ZEUS bounds on $\lambda_{eq1} \lambda_{\mu q2} / M_{LQ}^2$ in units of TeV ⁻² , for $\neq LQ$ types ed $\rightarrow \mu b$ 13 * $B \rightarrow r\bar{e} N$ ed $\rightarrow \mu s$ 32 * $B \rightarrow r\bar{e} N$ eb $\rightarrow \mu s$ 32 * $B \rightarrow r\bar{e} N$ $B \rightarrow r\bar{e} N$ better constrained by $\frac{D}{d} \lambda_{11} e^{-\frac{1}{d}}$ eb $\rightarrow \mu s$ 32 * $B \rightarrow r\bar{e} N$ $B \rightarrow r\bar{e} N$ Several examples where DIS constraints			le analysis	ti-variab	unu'	И Г		q1
q1multi-variable analysis10 $M < \sqrt{s}$ Constraints obtained on $(M_{LQ}, \lambda_e, \lambda_{\mu(t)})$ for $M < \sqrt{s}$ for $M < \sqrt{s}$ For very large masses : bounds on $\lambda_{eq} \lambda_{lq} / M^2 L_Q$ ZEUS, hep-ex/0501070To very large masses : bounds on $\lambda_{eq} \lambda_{lq} / M^2 L_Q$ ZEUS, hep-ex/0501070ZEUS bounds on $\lambda_{eq1} \lambda_{\mu q2} / M_{LQ}^2$ Better constrained byLQ μ 13 s 0.4 0.8 0.7 $d \rightarrow \mu b$ 13 s 0.4 0.4 0.4 $d \rightarrow \mu b$ 13 s 0.4 0.4 0.4 $d \rightarrow \mu b$ 13 s 14 0.4 0.4 $eb \rightarrow \mu s$ 32 s 14 0.4 0.4 $b \rightarrow \mu s$ 32 s 14 0.4 0.4 $b \rightarrow \mu s$ 12 s 0.4 0.4 0.4 $b \rightarrow \mu s$ 12 s 0.4 0.4 0.4 $b \rightarrow \mu s$ 12 s 0.4 0.4 0.4 $b \rightarrow \mu s$ s s s s s $b \rightarrow \mu s$ 12 s 0.4 0.4 0.4 $b \rightarrow \mu s$ s s s s s) GeV		ction.	inal selec hased oi	42 @ f 1-id		5	
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $		e [a] • ZEUS e [±] p 94-0 ■ Background M 10 ² Background M	: from µµ and ely suppressed	kgrounds DIS large) Bac CC I	ک ۲, ۲	. (a -
$\begin{array}{c c c c c c c c c c c c c c c c c c c $		s,		. <u>v</u>	rocesse	o LFV p	nd to	Could lec
Could lead to LFV processes. Could lead to LFV processes. Processes. Could lead to LFV processes. Processes. Processes. CC DIS largely suppressed Processes. CC DIS largely suppressed Processes. CC DIS largely suppressed Processes. CC DIS largely suppressed T-id based on a multi-variable analysis Constraints obtained on (MLQ, λ_e , $\lambda_{\mu(\tau)}$) for M $< \sqrt{s}$ For very large masses : bounds on λ_{eq} , $\lambda_{\mu}(\tau)$) for M $< \sqrt{s}$ For very large masses : bounds on λ_{eq} , $\lambda_{\mu}(\tau)$) for M $< \sqrt{s}$ For very large masses : bounds on λ_{eq} , $\lambda_{\mu}(\tau)$ or M $< \sqrt{s}$ TEUS bounds on λ_{eq1} , $\lambda_{\mu Q} / M^2 LQ$ TEUS hep-ex/0501070 DI Ω $\Delta = 0.4$ Ω_{10} ,		erent generations.	^c ermions of diff	pling to f	-Qs cou	o have l	لح t	Possibilit
Possibility to have LQs coupling to fermions of different generations. Could lead to LFV processes. Could lead to LFV processes. Could lead to LFV processes. Could lead to LFV processes. Packgrounds from µµ and C DIS largely suppressed η_1 η_2 η_1 η_2 η_1 η_2 η_1 η_2 η_3 η_4 η_1 η_2 η_4 $\eta_$		ng LQs	vor Violati	n Fla	epto	_		



<		del" : related,	rmine its do not depend	PJ C36 (2004) 425 =M _{squark} u _L -type sum	e ⁺ MJ, e ⁻ MJ	elMJ		$\begin{array}{c cccc} 150 & 200 & 250 \\ M_{\rm souark} (GeV) \end{array}$
straints within the MSSM	15 : all squark types, all decay modes top $ ightarrow$ eq and stop $ ightarrow$ b χ^+	Consider first an "unconstrained mod sfermion & gaugino sectors are not re i.e. M ~ are free parameters	Scan on the parameters which deter the other masses & branchings : limit much on the MSSM parameters.	Because S/B is good in B1 M1, EP1 all channels.	Sev)	αe_{m} = αe_{m} = $270 - 280 \text{ GeV}$ = 10^{-1}	$\frac{1}{2}$ indice the indirect limits $\frac{1}{2}$ (c)	150
Cons	- H1, EPJ C36 (2004) 42 - ZEUS Prelim : stop in s	ZEUS (prel.) 99-00 e ⁺ p 100 GeV < M2 < 300 GeV	10 ⁻¹ δ ¹³¹ (APV)	10 ⁻² Excluded in part of SUSY parame	M _{stop} (• For λ' = 0.3 (λ'^2 / 4π = squarks can be ruled of	 HERA's sensitivity exterior from low energy ext for 	E. Perez

hin mSUGRA	5 parameters. comparing to other experiments.	For intermediate m _o , excluded domain not ruled out by LEPII	DO bound (Run I) : moderate, again because of large branching in E _{T,miss} + jets	NB: Tevatron RunI : different frameworks, but typical sensitivities ≈ 130 GeV. RunII: ≈ 200-250 GeV	Discovery potential for RpV stop at HERA with larger L!	0 07/10/05
Constraints with	nore constrained SUSY model, with only 5 terpretation within mSUGRA useful when	80 tan $\beta = 6$ $\lambda'_{1j1} = 0.3$ 60 (b) (b) (b) (b) (b) (c)	$40 = 10^{-1.1} \text{ H}$		20 - DU KUN 1 not allowed	^V 0 50 100 150 200 250 300 350 1, EPJ C36 (2004) 425 m ₀ (GeV)



A new scale of matter ? I+V re signature : direct observation of excited s f_{α} (chiral) magnetic coupling \rightarrow (GeV $\Lambda \approx$ compositeness scale Relative strength of γ , Z and g couplings \rightarrow f, f', f _s but and g couplings \rightarrow f, f', f _s but Pair production of f* in e ⁺ e ⁻ and pp ; single product $\stackrel{He}{}_{\alpha}$ (chiral) in e ⁺ e ⁻ and pp ; single product $\stackrel{He}{}_{\alpha}$ (compton of f* in e ⁺ e ⁻ and pp ; single product $\stackrel{He}{}_{\alpha}$ (compton at HERA $\stackrel{He}{}_{\alpha}$ (compton at HERA H	esonances ?	rates) -1 giwara et al, ZPC 29 (1985) 115. Idjema et al, ZPC 57 (1993) 425.	ion depends on coupling	ed excited fermions, (ey) resonance NB: e ⁻ data very sensitive to v [*] . HERA II will improve a lot w.r.t our HERA I results. (7000
	A new scale of matter ? I+V re	Inambiguous signature : direct observation of excited st f f f (chiral) magnetic coupling \rightarrow (GeV f^* $\Lambda \approx$ compositeness scale Relative strength of γ , Z and g couplings \rightarrow f, f', f _s Bound	\rightarrow Pair production of f [*] in e [*] e ⁻ and pp ; single production	$\begin{array}{c c} & \text{Eus} & \text{for singly product} \\ & \text{ELPHI + OPAL, Direct (Prelim)} \\ & \text{DELPHI + OPAL, Direct (Prelim)} \\ & \text{DERMIN} \\ & DER$

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18

07/10/05





H1 dedicated top search (HERA I)

Excess observed in H1 compatible with top production ? Make use of :

- the lepton charge : valence quark \rightarrow u \rightarrow t (not $\overline{1}$) $\rightarrow \oplus$ H1 Collab., EPJ. C33 (2004) 9 **ELECTRON CHANNEL**
 - angular distributions : cos θ_{W} (cf W helicity in top decays) - exploit the large expected ${\sf P}_{{\sf T},{\sf jet}}$ & M $_{\sf lvb}$

All SM processes

Top MC

5tn∋v∃ 5

H1 Data

5 (3e + 2µ) of the HERA-I "isolated lepton evts" appear top-like ! SM expectation = 1.31 \pm 0.22

Large γp + DIS multijet background (W subleading). Search also carried out in the 3-jet channel.

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	leen			s well	59	ie had. that	Susy	ج ج	07/10/05
	H1 excess not s	ine ⁻ p!	3.4 σ effect	n at the Tevatron a	H1, PLB 599 (2004) 15	Jnlikely because of th thannel (jij + P _{T,miss} in tase, good S/B)	b RpV	X+X T LSP =	fy the excess !!
p.	Combined e & μ	2 / 1.8 ± 0.3	15 / 4.6 ± 0.8	to appear <mark>soo</mark> l	۲ ۲ ۲		+ \$ •	λ' ₁₃₁	es. First clari
obs. / ex	µ channel	0 / 0.9 ± 0.2	6 / 2 .3 ± 0.4	f NP, likely	+o	2 V ¹³¹	, cuila	e. fields ch P _{T,miss} in	e possibiliti
: pb-1	e channel	2 / 0.9 ± 0.2	<mark>9</mark> / 2.3 ± 0.4	[:] luctie. But i	ers 22	х ^с \	Two DnV cou	with 3 rd gen (Not too mud that case)	There ar
11 Preliminary, 211	P _T X > 25 GeV	Electrons, 98-05 53 pb-1	Positrons, 94-04 158 pb-1	ight be a 3.40 f	ssible explanati	"LQ"	, μ,d λ'133	+ 2 2 4 - 0 + 0 -	$\sqrt{\frac{1}{\lambda^{313}}}$
	H1 Preliminary, 211 pb-1 obs. / exp.	H1 Preliminary, 211 pb-1 obs. / exp. P _T ^X > 25 GeV e channel μ channel Combined e & μ H1 excess not seen	H1 Preliminary, 211 pb-1obs. / exp. $P_T^X > 25 \text{ GeV}$ e channel μ channel μ channel $R1 \text{ excess not seen}$ Flectrons, 98-05 $2 / 0.9 \pm 0.2$ $0 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ $1 \text{ in e} \text{ p} \text{ l}$	H1 Preliminary, 211 pb-1obs. / exp. $P_T^X > 25 \text{ GeV}$ e channel μ channelCombined e & μ $P_T^X > 25 \text{ GeV}$ e channel μ channelCombined e & μ Electrons, 98-05 $2 / 0.9 \pm 0.2$ $0 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ 53 pb-1 $2 / 0.9 \pm 0.2$ $0 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ Positrons, 94-04 $9 / 2.3 \pm 0.4$ $6 / 2.3 \pm 0.4$ $15 / 4.6 \pm 0.8$ 158 pb-1 $9 / 2.3 \pm 0.4$ $15 / 4.6 \pm 0.8$ 3.4σ effect	H1 Preliminary, 211 pb-1obs. / exp. $p_T^X > 25 \text{ GeV}$ $e \text{ channel}$ $\mu \text{ channel}$ $combined e \& \mu$ $p_T^X > 25 \text{ GeV}$ $e \text{ channel}$ $\mu \text{ channel}$ $\rho \text{ combined e } \& \mu$ $E ectrons, 98-05$ $2 / 0.9 \pm 0.2$ $0 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ 53 pb-1 $p - 23 \pm 0.4$ $6 / 2.3 \pm 0.4$ $15 / 4.6 \pm 0.8$ Positrons, 94-04 $9 / 2.3 \pm 0.4$ $6 / 2.3 \pm 0.4$ $15 / 4.6 \pm 0.8$ 158 pb-1 $9 / 2.3 \pm 0.4$ $15 / 4.6 \pm 0.8$ Might be a 3.4 \sigma fluctie. But if NP, likely to appear soon at the Tevatron as well	H1 Preliminary, 211 pb-1obs. / exp. $P1$ Preliminary, 211 pb-1obs. / exp. $P1$ Preliminary, 211 pb-1 $p2$. / $p3$. / $p3$ $P1$ Preliminary, 211 pb-1 $p3$. / $p3$ $P2$ Provins, 98-05 $2 / 0.9 \pm 0.2$ $0 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ $2 / 0.9 \pm 0.2$ $2 / 1.8 \pm 0.3$ Positrons, 94-04 $9 / 2.3 \pm 0.4$ $(5 / 2.3 \pm 0.4)$ $(5 / 4.6 \pm 0.8)$ Positrons, 94-04 $9 / 2.3 \pm 0.4$ $(5 / 2.3 \pm 0.4)$ $(5 / 4.6 \pm 0.8)$ Positrons, 94-04 $9 / 2.3 \pm 0.4$ $(5 / 4.6 \pm 0.8)$ $(3.4 \sigma \text{ effect})$ Might be a 3.4 \sigma fluctie. But if NP, likely to appear soon at the Tevatron as well $(7 - \sqrt{6})$ Possible explanations ?? $(7 - \sqrt{6})$ $(1, PLB 599 (2004) 159)$	H1 Preliminary, 211 pb-1obs. / exp. $P_{T}^{X} > 25 6eV$ e channel μ channel <t< th=""><th>H1 Preliminary, 211 pb-1 obs. / exp. Pr.Y. 25 GeV e channel μ channel α mined e & μ Electrons, 98-05 2 / 0.9 \pm 0.2 0 / 0.9 \pm 0.2 2 / 1.8 \pm 0.3 53 pb-1 2 / 0.9 \pm 0.2 2 / 1.8 \pm 0.3 Bositrons, 94-04 9 / 2.3 \pm 0.4 6 / 2.3 \pm 0.4 15 / 4.6 \pm 0.8 Might be a 3.4 of fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible expl</th><th>H1 Preliminary, 211 pb-1 obs. / exp. Pr7 × 25 6 eV e channel μ channel μ channel α in α for the contined e 4μ Electrons, 98-05 $2/0.9 \pm 0.2$ $0/0.9 \pm 0.2$ $2/1.8 \pm 0.3$ Electrons, 98-05 $2/0.9 \pm 0.2$ $0/0.9 \pm 0.2$ $2/1.8 \pm 0.3$ Positrons, 94-04 $9/2.3 \pm 0.4$ $6/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Positrons, 94-04 $9/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Might be a 3.45 fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible explanations ?? Positrons, 94-04 $0/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Might be a 3.45 fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible explanatint ?? Possible explanations ?? Possible explanatio</th></t<>	H1 Preliminary, 211 pb-1 obs. / exp. Pr.Y. 25 GeV e channel μ channel α mined e & μ Electrons, 98-05 2 / 0.9 \pm 0.2 0 / 0.9 \pm 0.2 2 / 1.8 \pm 0.3 53 pb-1 2 / 0.9 \pm 0.2 2 / 1.8 \pm 0.3 Bositrons, 94-04 9 / 2.3 \pm 0.4 6 / 2.3 \pm 0.4 15 / 4.6 \pm 0.8 Might be a 3.4 of fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible expl	H1 Preliminary, 211 pb-1 obs. / exp. Pr7 × 25 6 eV e channel μ channel μ channel α in α for the contined e 4μ Electrons, 98-05 $2/0.9 \pm 0.2$ $0/0.9 \pm 0.2$ $2/1.8 \pm 0.3$ Electrons, 98-05 $2/0.9 \pm 0.2$ $0/0.9 \pm 0.2$ $2/1.8 \pm 0.3$ Positrons, 94-04 $9/2.3 \pm 0.4$ $6/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Positrons, 94-04 $9/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Might be a 3.45 fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible explanations ?? Positrons, 94-04 $0/2.3 \pm 0.4$ $15/4.6 \pm 0.8$ Might be a 3.45 fluctie. But if NP, likely to appear soon at the Tevatron as well Possible explanations ?? Possible explanatint ?? Possible explanations ?? Possible explanatio



	Anomalo	us multilepton	production	C. L
If anomalor	us W productior	n: what about anomalou	دد Z s	e^+ e^-
Events with via 77. Cross	$n \ge 2$ leptons in t s-section ~ 1 pb	final state. Mainly prod for central I, Pt > ~ 5	duced et	
-		- - (H1 94-00 data	obs. / exp.
HI data ya high M ₁₂ =	t-UU : excess of mass of two hid	· ze+3e events at ghest P _T e	expt selection	H1 (115 pb ⁻¹)
No such ex	kcess seen in ZE	EUS HERA-I data	2e, M ₁₂ > 100 GeV	3 / 0.30 ± 0.04
			3e, M ₁₂ > 100 GeV	$3 / 0.23 \pm 0.04$
H1 analysis	extended to in	clude 03-05 data :	H1, EPJ C31 (2003)	17
209 pb ⁻ 1 ^{0*} 탈	-1 H1 Prelim., HERA I+II	Extended to oth	er 21 & 31 topolog	ies :
Even	t	Νοω εε, μμ, εμ,	eee, eµµ are cons	sidered
₽		 no new 2e / 3e evt c 	11 M ₁₂ > 100 GeV	
<u></u>		(but one high mass a • one eμμ evt at M _{μμ}	se event) > 100 GeV, one at	M _{eµ} > 100 GeV
10 ⁻ 10 + 31		Altogether, at 2P	T > 100 GeV :	
10 ⁻ - 10 ⁻ - 10 0 20 40 60 E. Perez	0 80 100 120 140 160 18 2 P _T (GeV)	N _{obs} = 4, N _{exp} =	0.81 ± 0.14	07/10/05

"Signature Based" Searches for NP

(Quasi) "model-independent" search for new physics :

- definition of objects (e, μ, ν, γ, jet, W, Z, ...)
 look at data vs SM in all "channels" with > 1 object
- ullet in each channel, find the part of eta space with largest deviation (e.g. in M, Σ p $_{ op}$)
 - · quantify the agreement using "Gedanken" (Mock, MC) expts





Pionnered by DZero with D0, PRD64, 012004 (2001) the full Run I sample



26