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# Prompt Photons

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Introduction

Recent results in

$\gamma p, ep$

$pp, pN$

Pb Pb

Au Au

Some Questions

(apologies for data not covered)

# Light played always a central role in attempts to understand early states



Fastentuch, Zittau, 1472 AD

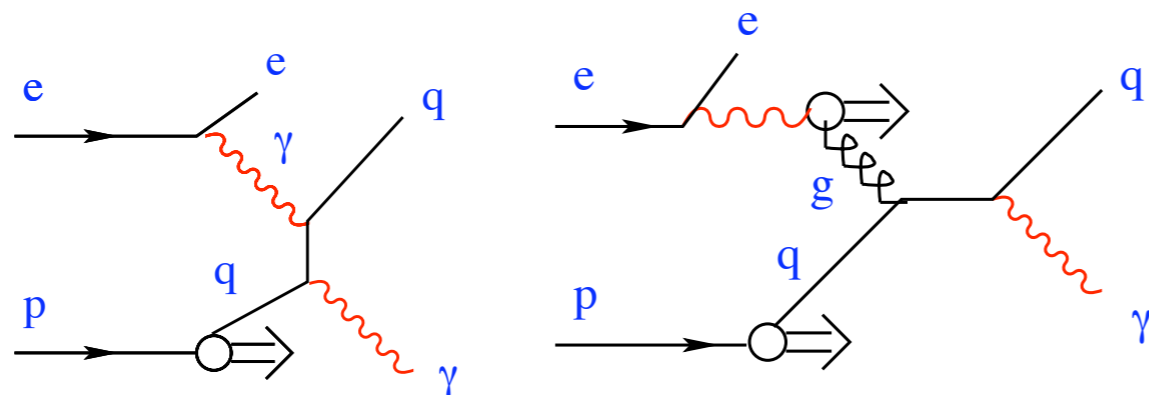
- Early reports on world creation
- Seeing the universe back from now to some  $10^5$  years after big bang
- Observing parton patterns through fire balls of nucleus-nucleus interactions
- Studying hard partonic interactions directly

(consider the latter two)

Usually photons are called prompt (direct) if coupling to interacting partons

in contrast to photons from hadron decays or radiation from leptons

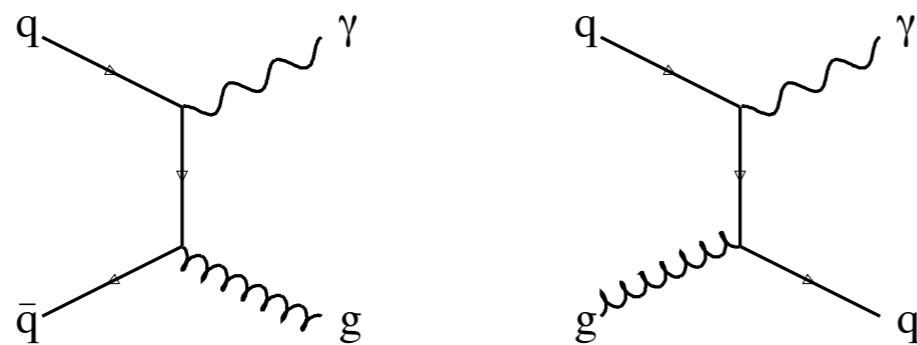
ep LO



examples of direct and resolved photon interactions

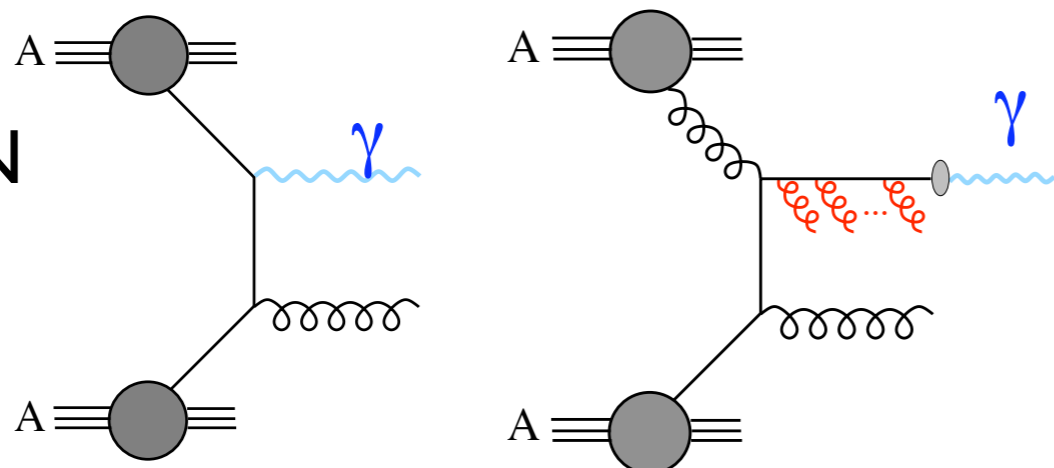
sensitive to gluon content of photon

hh LO



and of proton

NN



signal of hard interactions

but in all reactions contributions of non perturbative fragmentation processes

## Interest in prompt photons, because

- more directly related to partonic interactions than jets
- sensitive to gluon content of proton (and resolved photon)
- important background for searches at LHC (e.g. Higgs  $\rightarrow \gamma\gamma$ )
- they help in nuclear interactions to disentangle different effects (initial/final states, QGP, hadron gas, ...), as not strongly interacting

### drawbacks

— small cross sections

— background from neutral hadrons ( $\pi^0, \eta\dots$ )

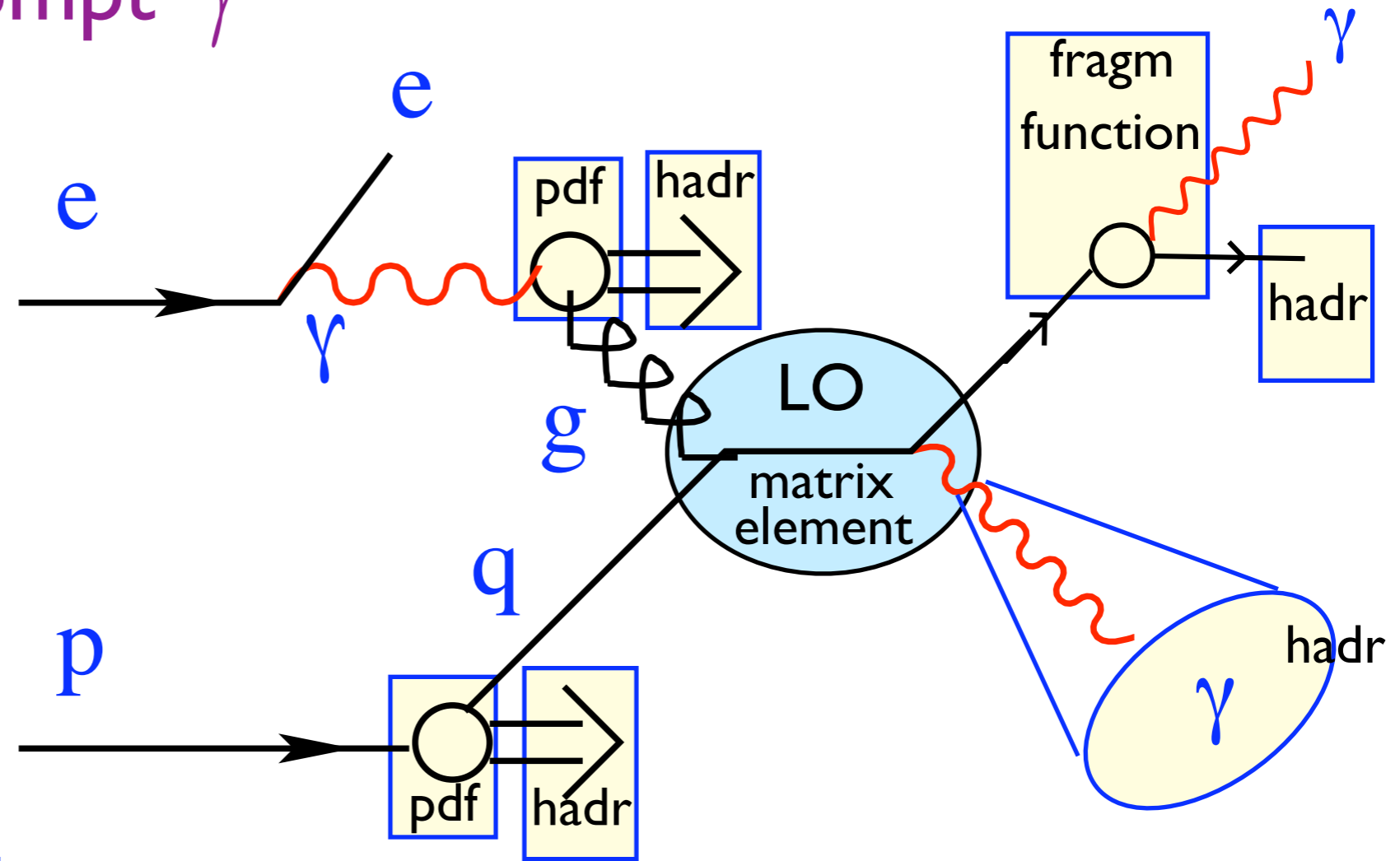
➔ needs careful subtraction of  $\pi^0, \eta\dots$

and/or sophisticated shower shape analyses

# Example of prompt $\gamma$ diagram

various calculations with NLO pQCD matrix elements and collinear pdfs ( $\gamma p$ ,  $e p$ ,  $\gamma\gamma$ ,  $hh$ )

improvements in  $hh$  by soft gluon resummations



non-perturbative elements

most recent analysis

Lipatov, Zotov(2005) :  $k_T$  factorisation for  $\gamma p$ ,  $pp$   
 (earlier work: Kimber, Martin, Ryskin (2000))

improved kinematics, off-shell matrix elements

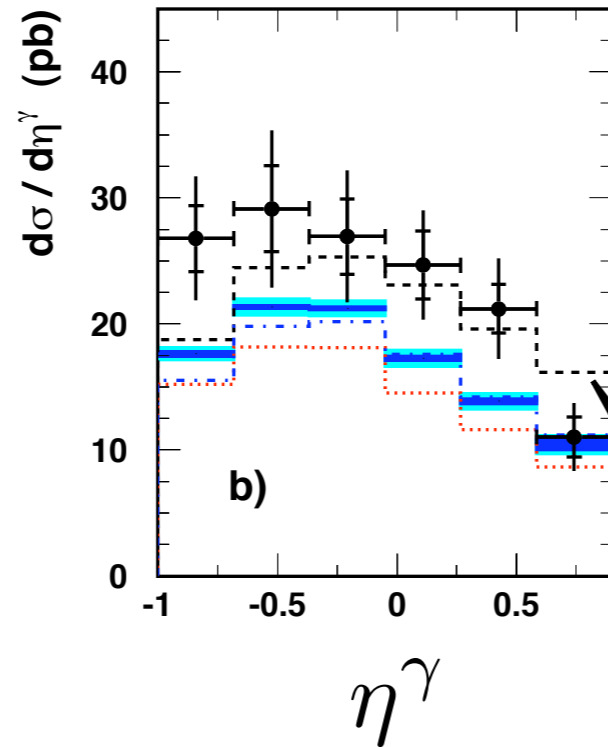
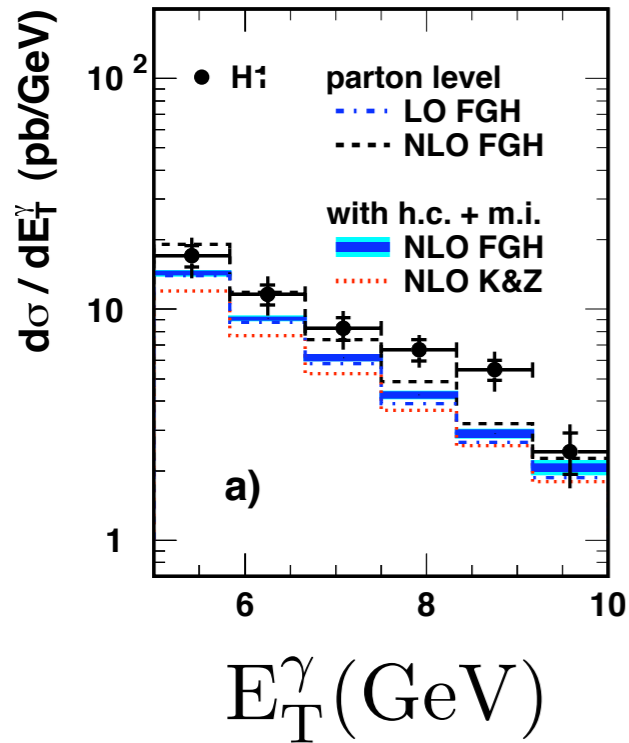
LO results close to collinear NLO !

# $E_T$ and pseudorapidity distributions in $\gamma p$ , HERA

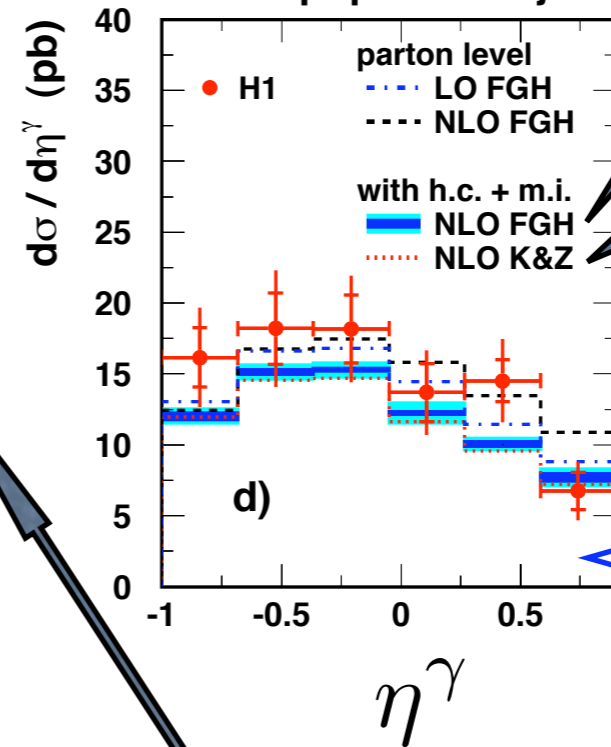
$$p \rightarrow \leftarrow \gamma$$

$$\sqrt{s_{\gamma p}} \sim 200 \text{ GeV}$$

Inclusive prompt photon



Prompt photon + jet



NLO pQCD

Fontanaz, Guillet, Heinrich

Krawczyk, Zembrzuski

description improved, if jet required

NLO and models PYTHIA, HERWIG describe shapes, but low in normalisation ~30...40%

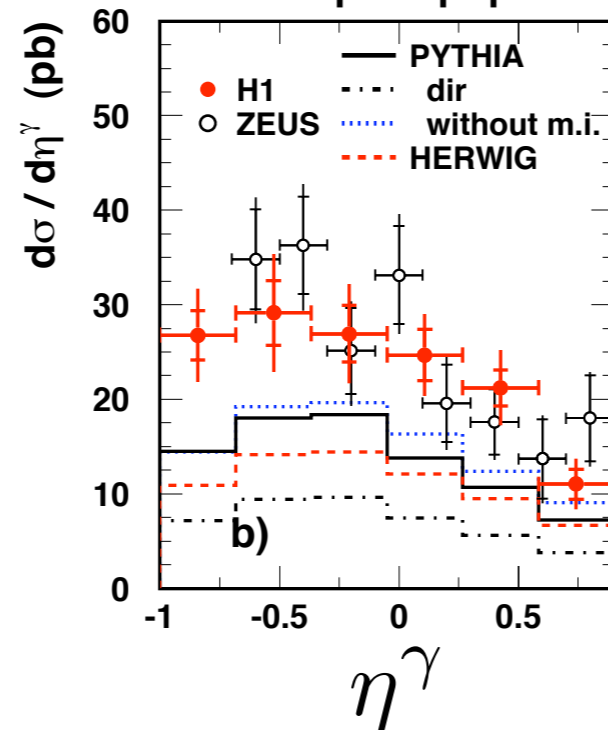
PYTHIA similarly low in  $\gamma\gamma$  (OPAL)

note :

NLO before corrections for hadronisation is closer to the data

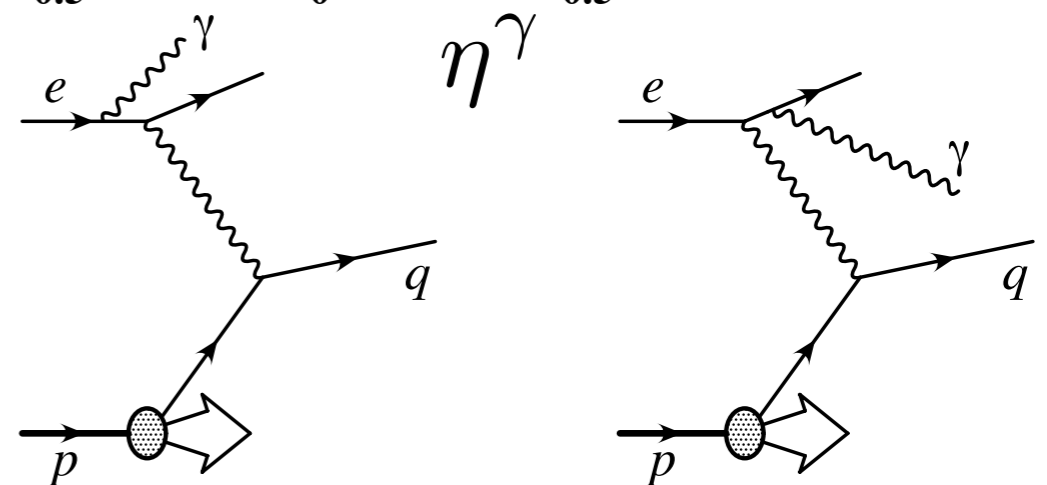
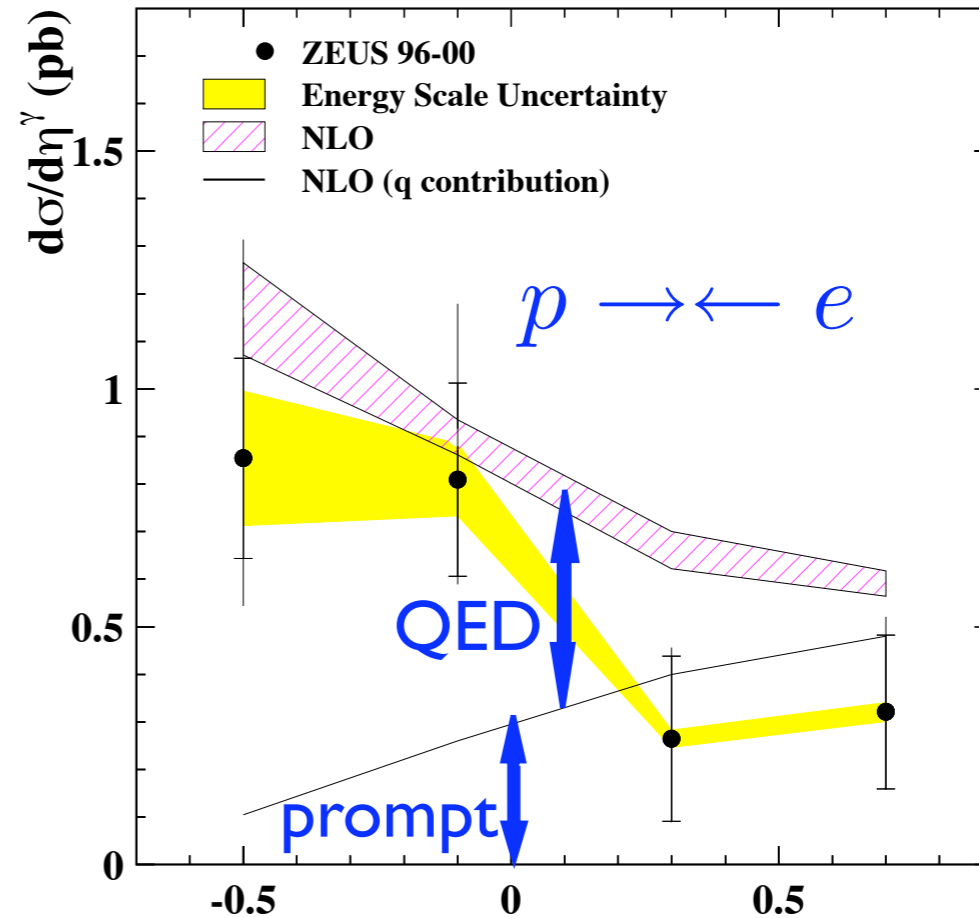
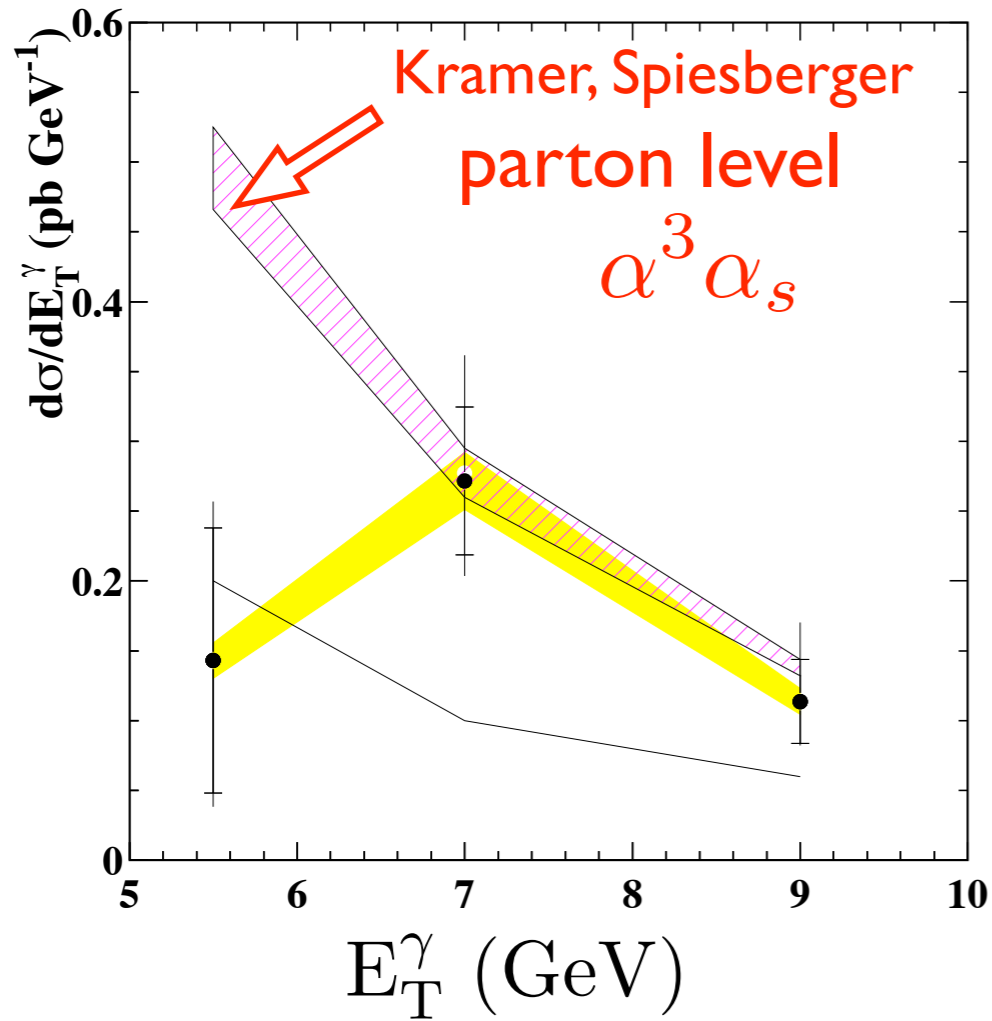
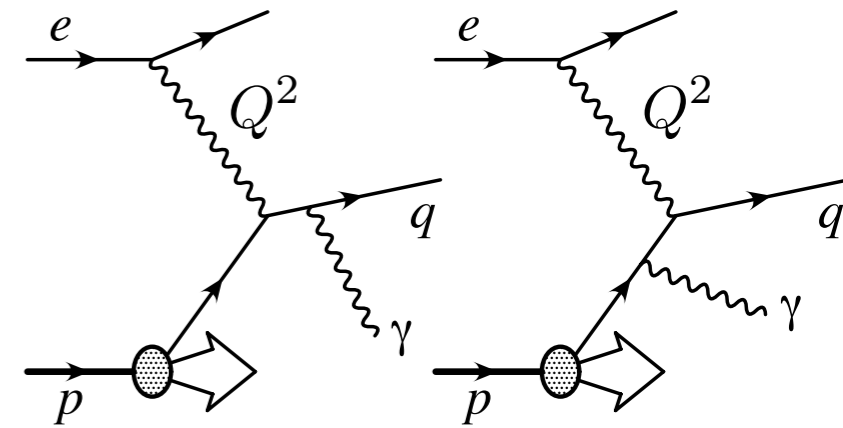
( H1 corrects with PYTHIA and HERWIG )

Inclusive prompt photon



# First results in DIS

$$Q^2 > 35 \text{ GeV}^2$$

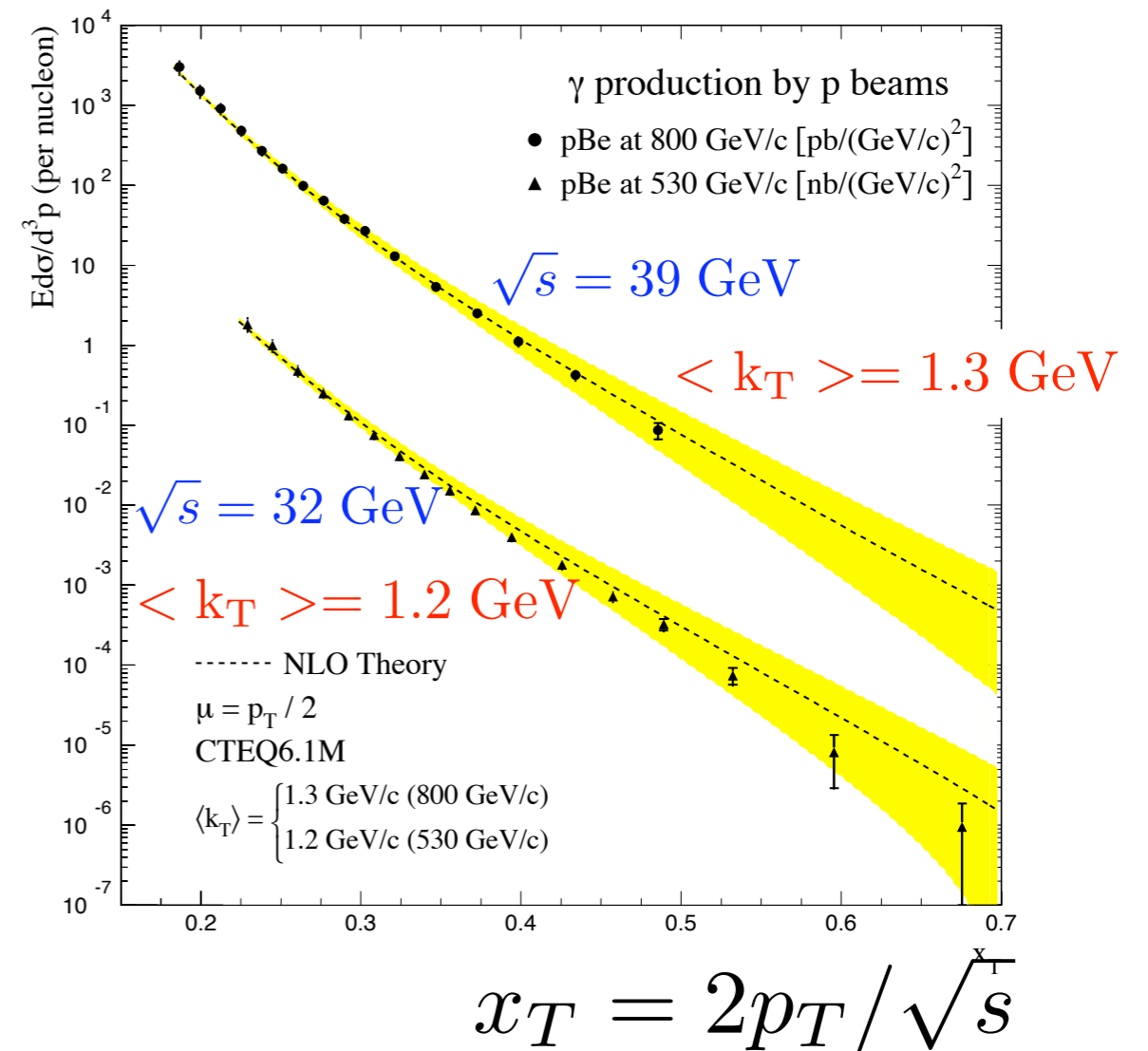
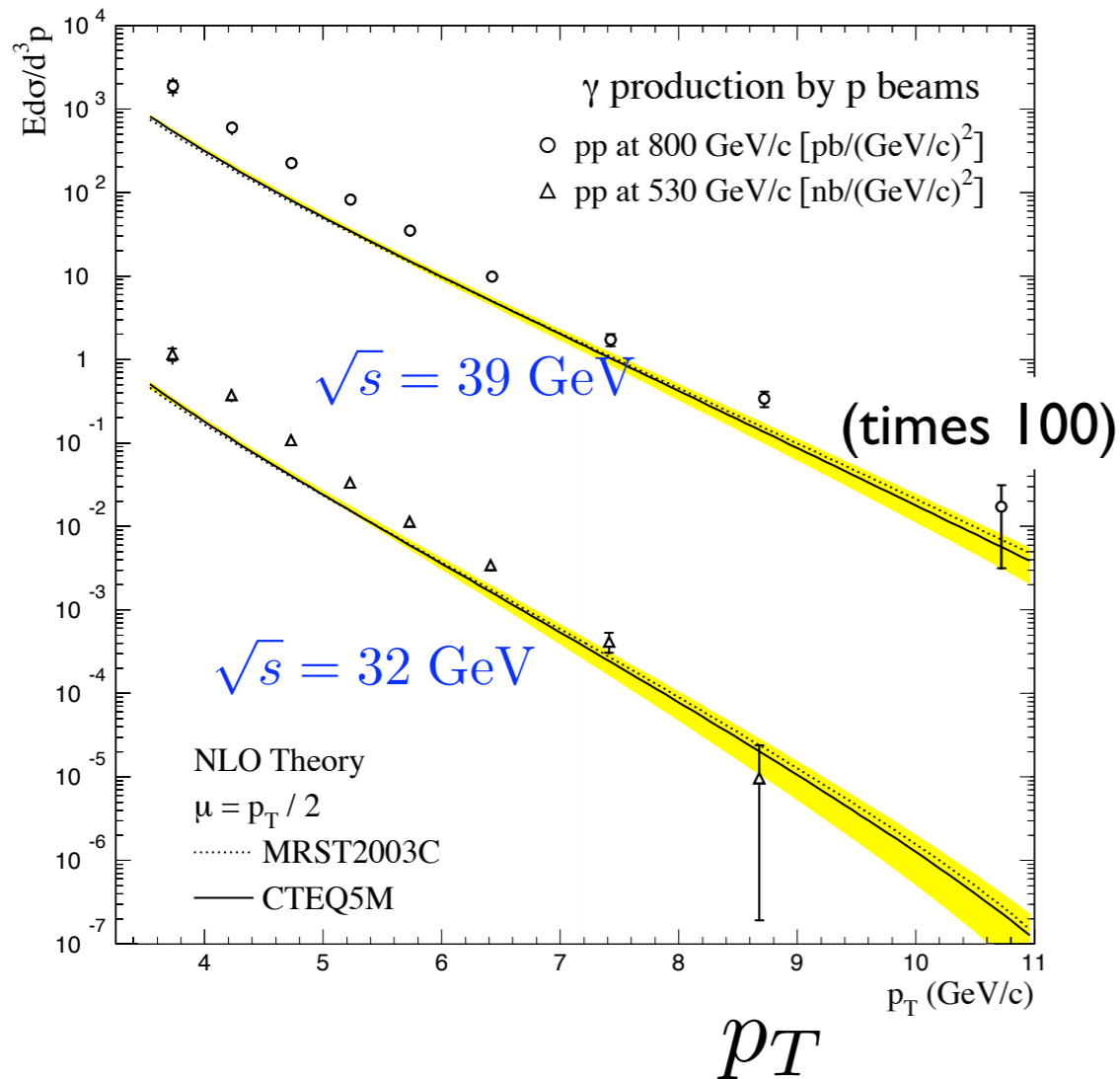


substantial contribution of QED  
uncertainties still large

data consistent with NLO

$$pN \rightarrow \gamma X$$

notoriously difficult to describe low energy data  
example of E706 Phys Rev D70 (2004), final report



**NLO** (Aurenche et al., Berger, Quin) far below data (< 50%) at small  $p_T$   
**Agreement reached by ad hoc smearing of initial  $\langle k_T \rangle \gtrsim 1$  GeV**

improvements by resummations (Lai, Li, 1998, Laenen, Sterman, Vogelsang, 2000

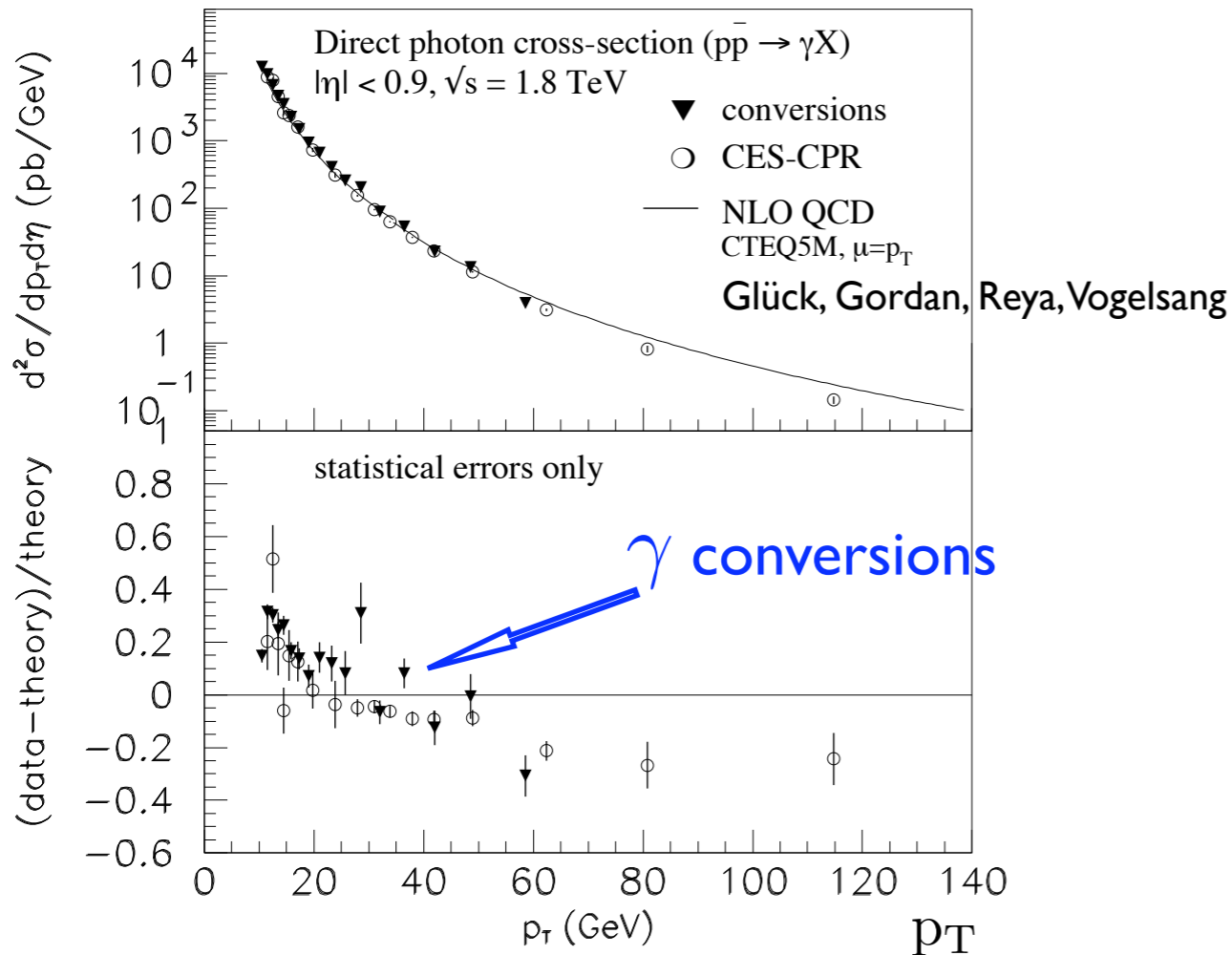
Kidonakis, Owens, 2003 ... Florian, Vogelsang, 2005, Sterman, Vogelsang 2005.....)



# $pp \rightarrow \gamma X$ , high energies

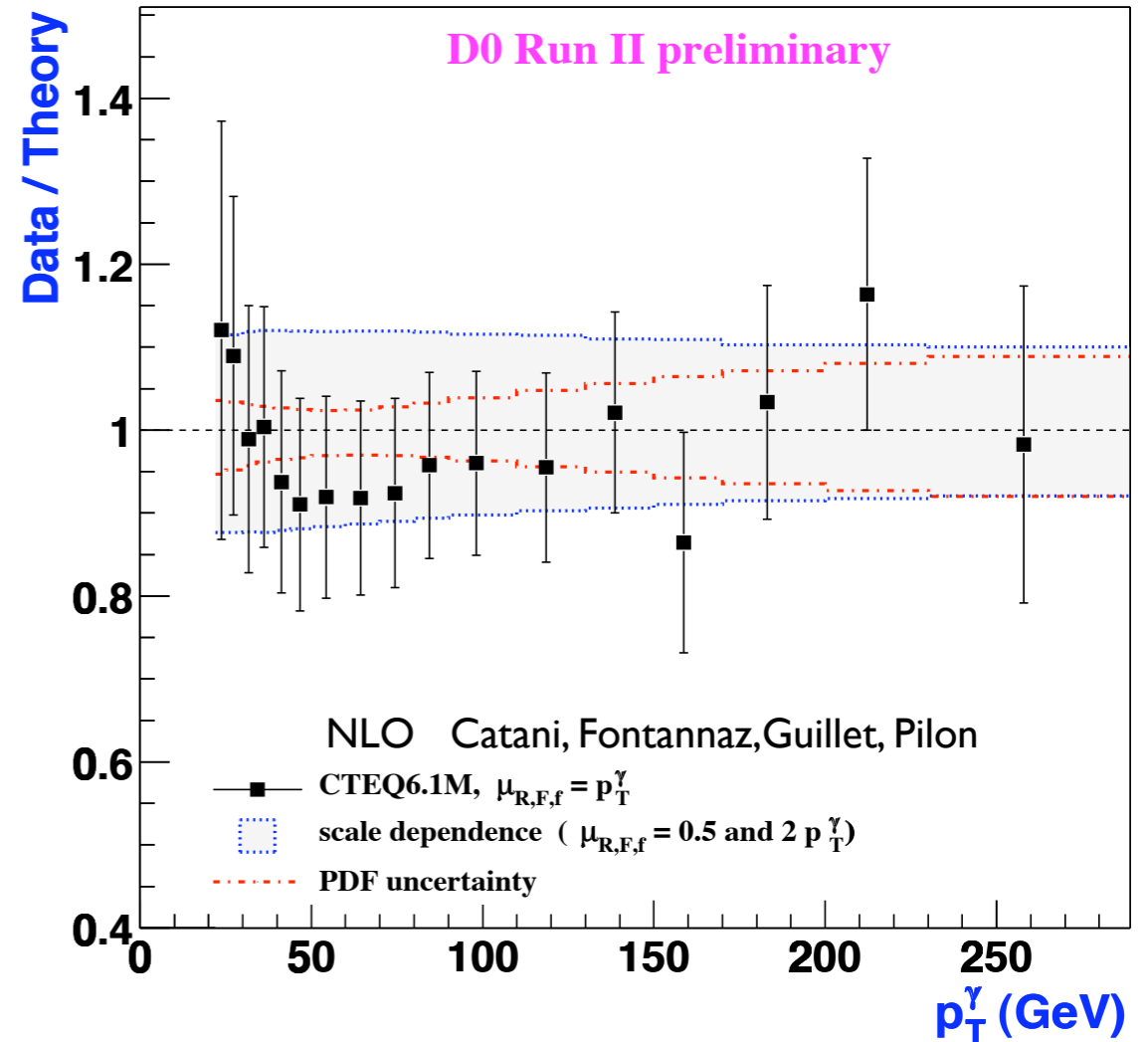
CDF

$\sqrt{s} = 1.8$  TeV



D0

$\sqrt{s} = 1.96$  TeV



in old and new CDF data  
 steeper  $E_T$  dependence than NLO

prel. D0 consistent with NLO

but in general, agreement better at high  $s$

less need for corrections like soft gluon resummations ?

# Nucleus-Nucleus Collisions

prompt photons may disentangle  
effects of nuclear medium

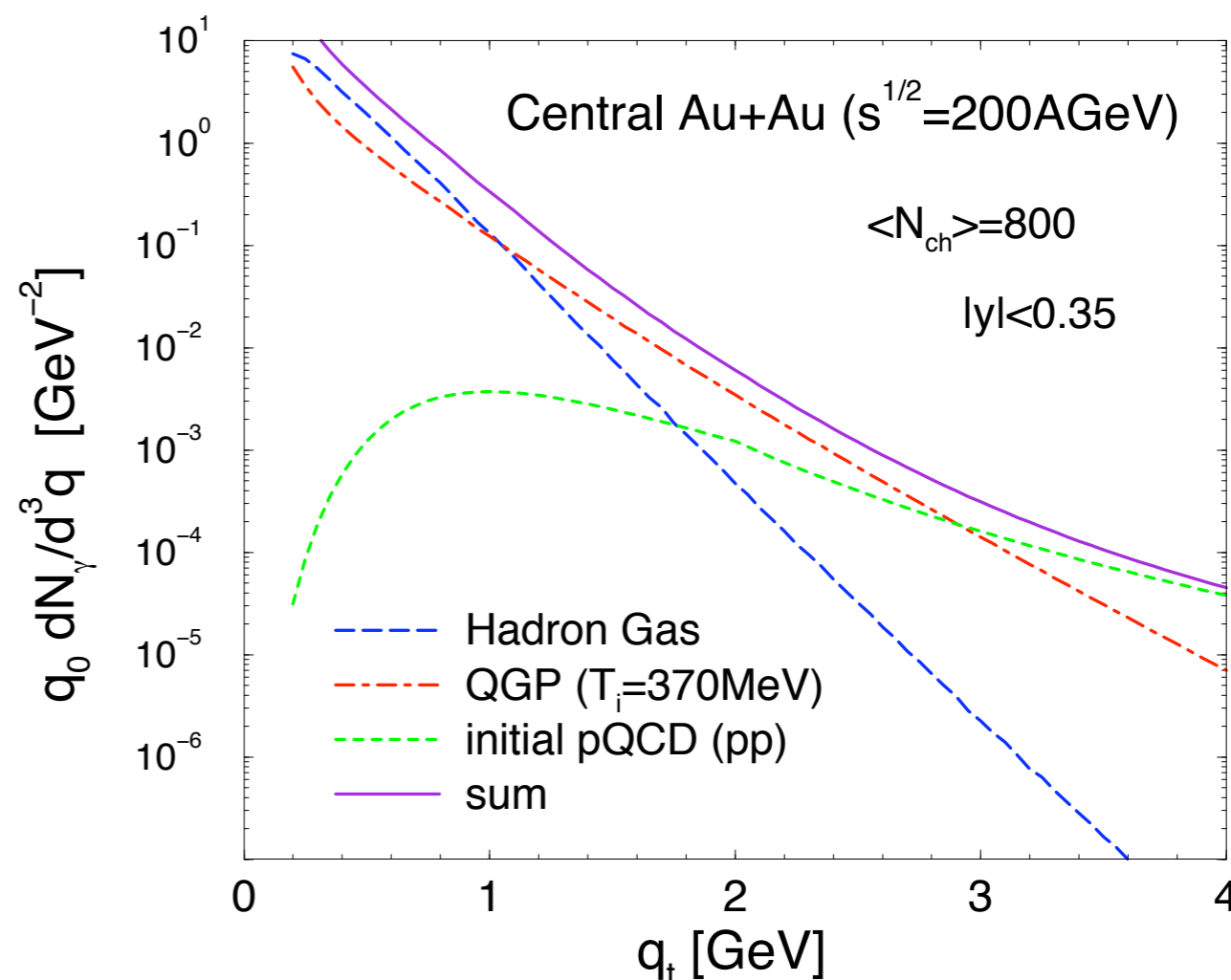
in particular  
thermal photons are expected  
from

QGP (quark-gluon-plasma)  
or Hadron Gas

expect sensitivity for QGP at medium  $p_T$

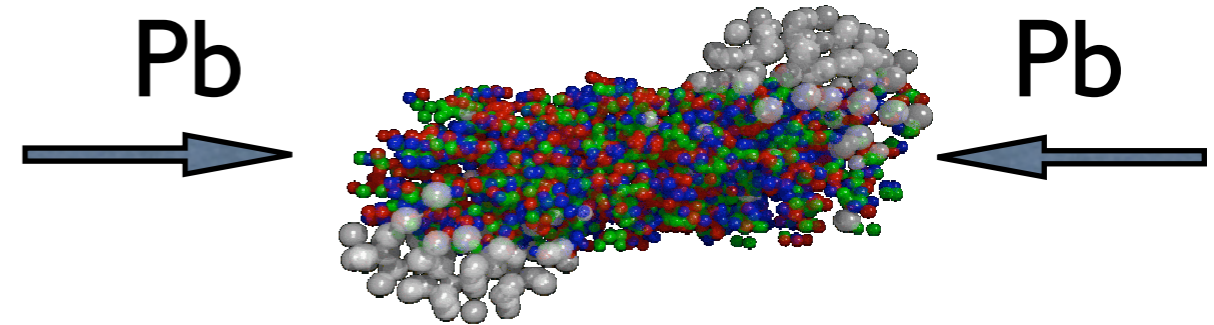
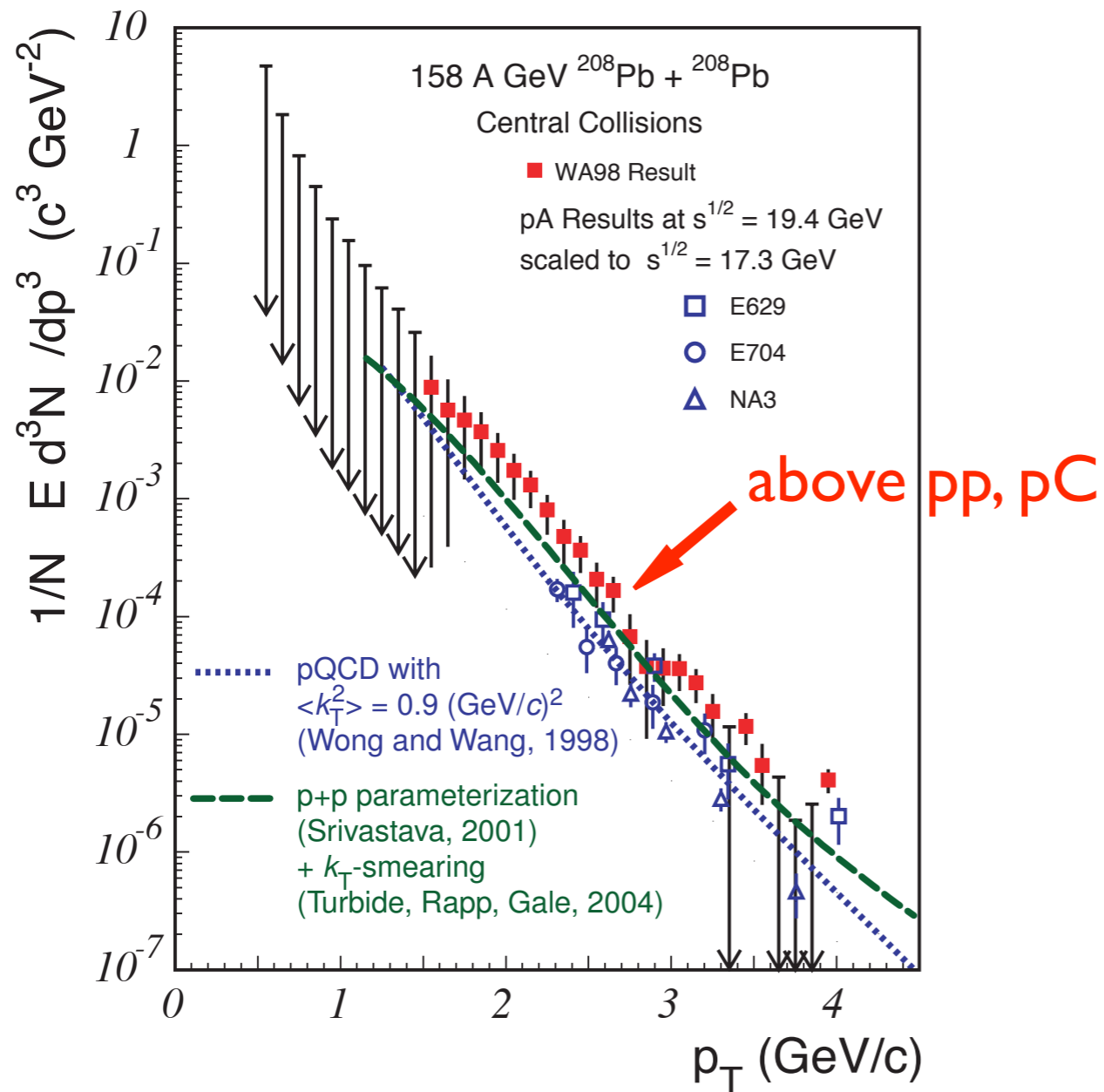
but pQCD (pp) dominates already at moderate  $p_T$

Turbide, Rapp, Gale

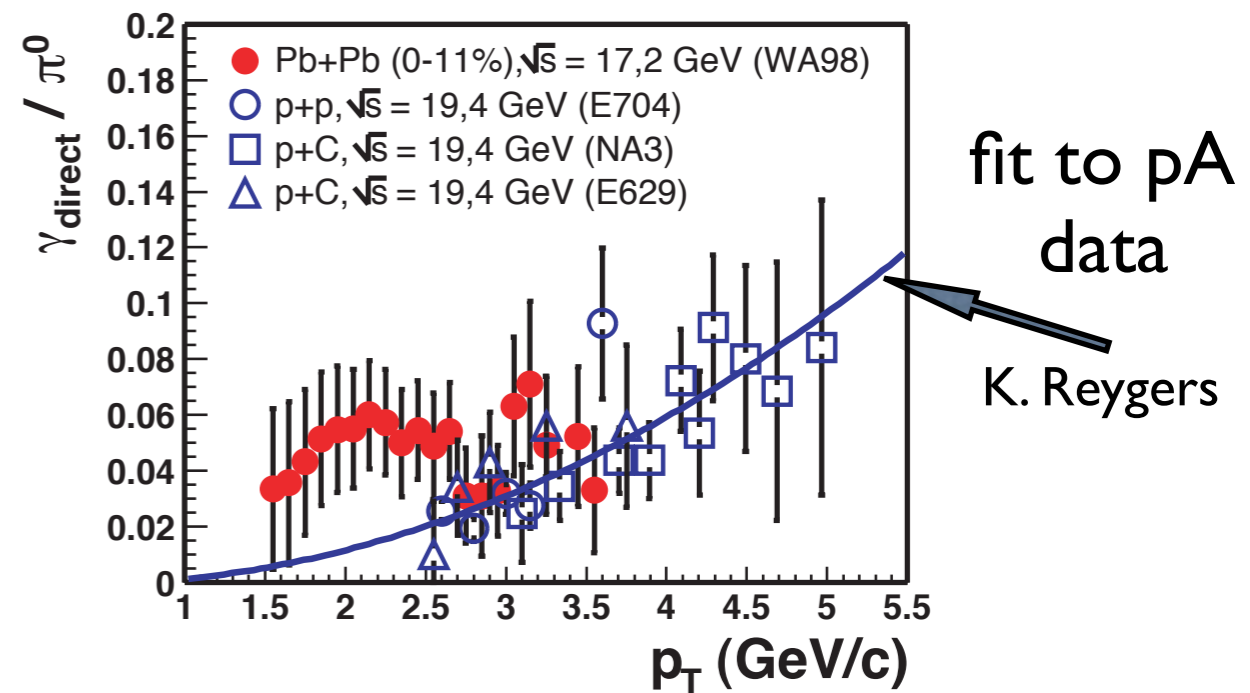


# Pb-Pb Collisions

CERN SPS, WA98  $\sqrt{s_{NN}} = 17.3 \text{ GeV}$



$$\gamma_{\text{dir}} / \pi^0$$

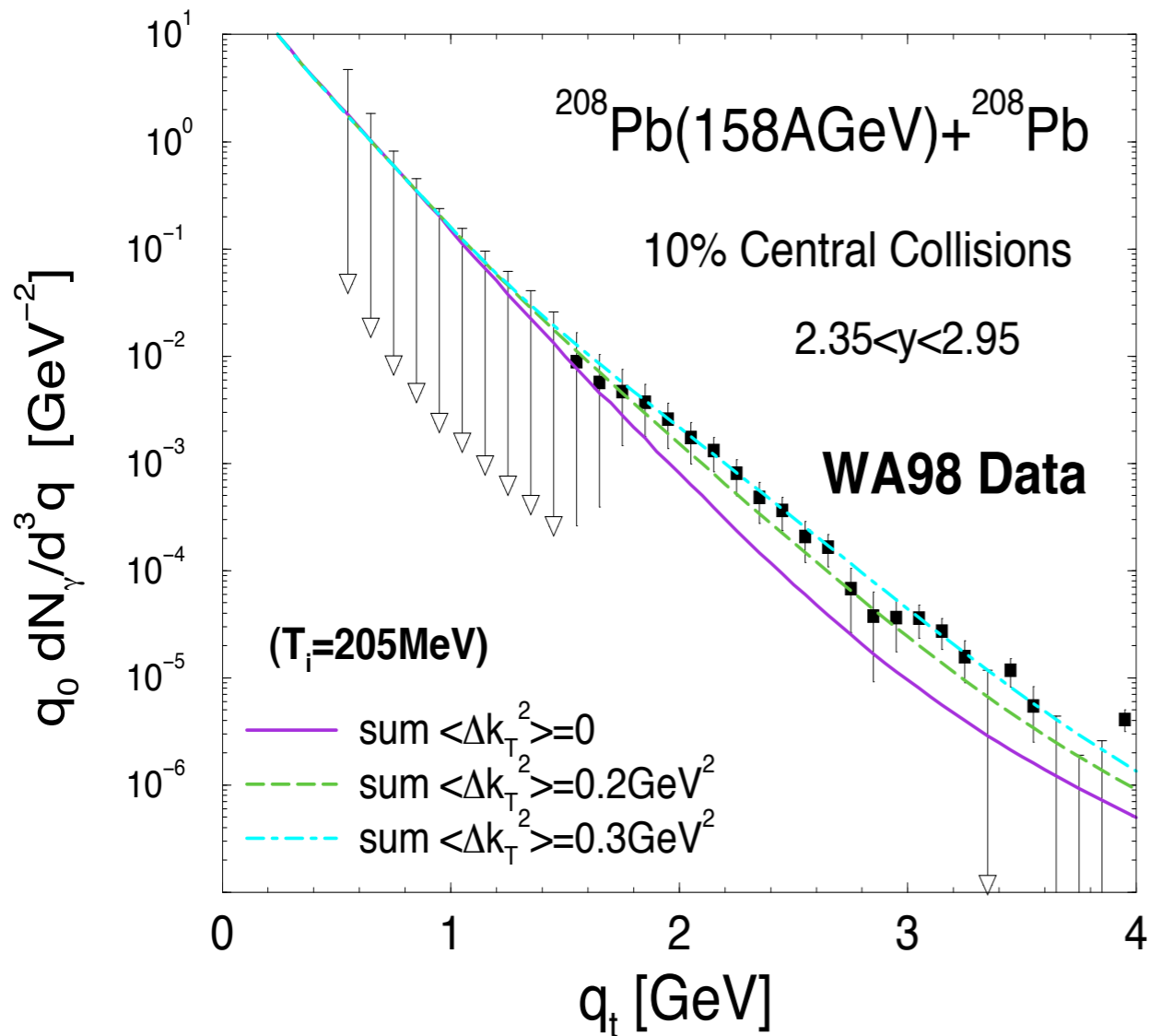


difficulties:

- large background at low  $p_T$
- no pA data at low  $p_T$
- no pPb data available

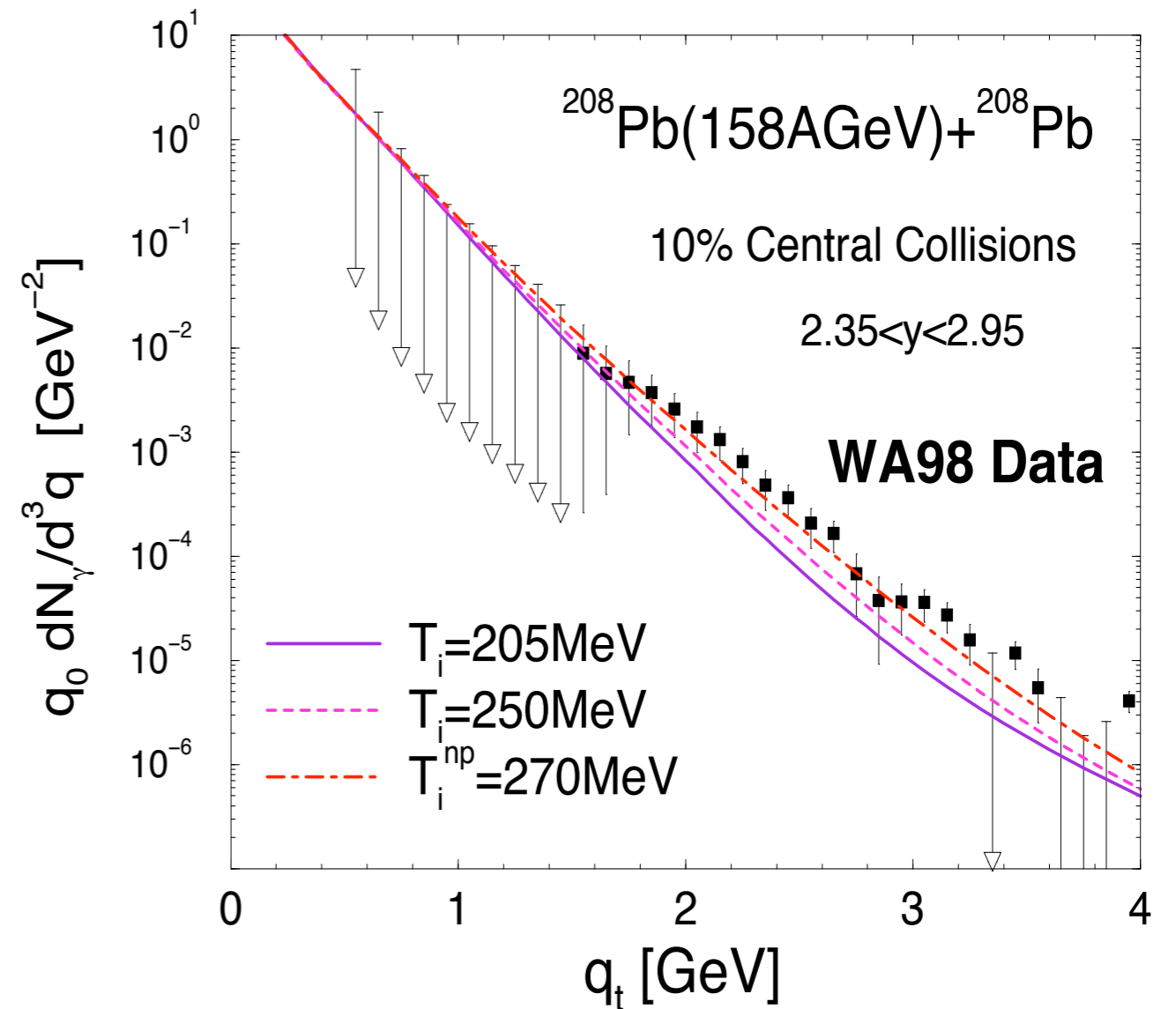
¿ QGP or hadron gas seen ?

# WA98 data as convolution of thermal photons of QGP and hadron gas and pQCD (Turbine, Rapp, Gale)



data described by

initial temperature 205 MeV  
and  $k_T$  broadening

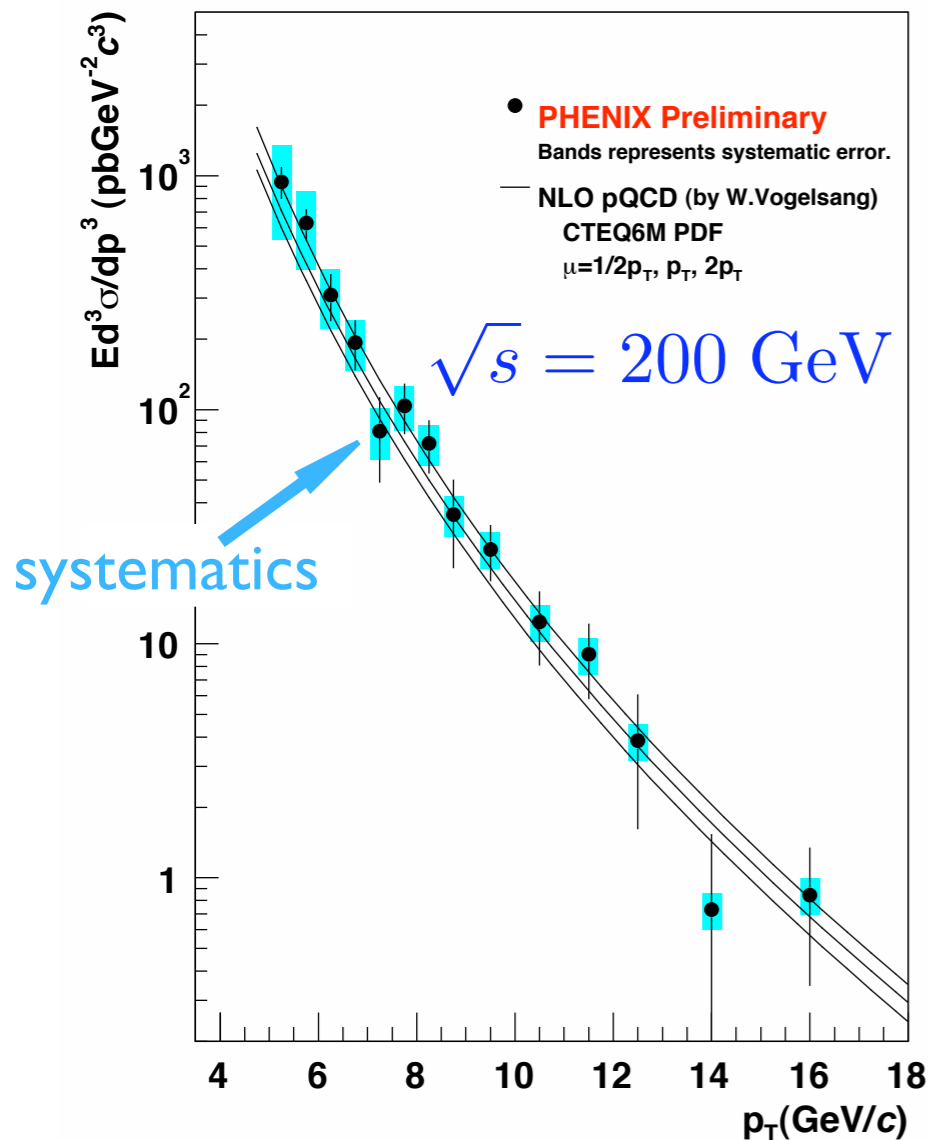


or

high initial temperature 270 MeV

# Results from RHIC

pp PHENIX



consistent with NLO

Gordon and Vogelsang

but what happens with heavy nuclei at these ultra-relativistic energies ?

we know

at central collisions

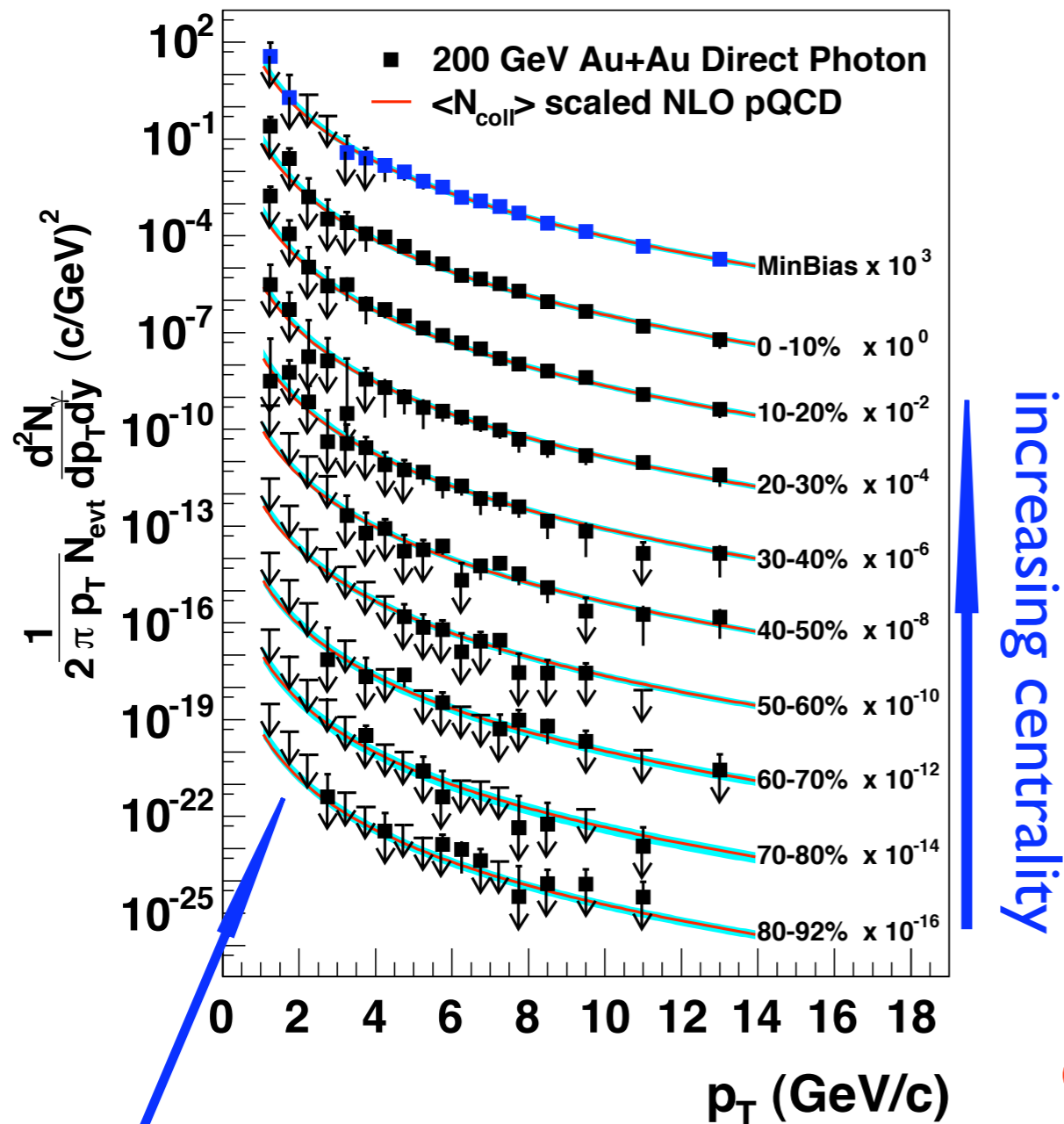
- high  $p_T$  hadrons are suppressed
- away side jets disappear at high  $p_T$  indicative of “jet quenching” due to parton energy loss in coloured medium

initial state effects

would suppress photons too

# Au+Au collisions

PHENIX  $\sqrt{s_{NN}} = 200$  GeV



NLO (Gordon, Vogelsang)

scaled from pp to  $\langle N_{\text{coll}} \rangle$  in AuAu

$N_{\text{coll}}$

number of inelastic NN collisions  $\sim$  nuclear overlap, deduced with Glauber theory from signals of beam counters and zero degree calorimeters

remarkable consistency !

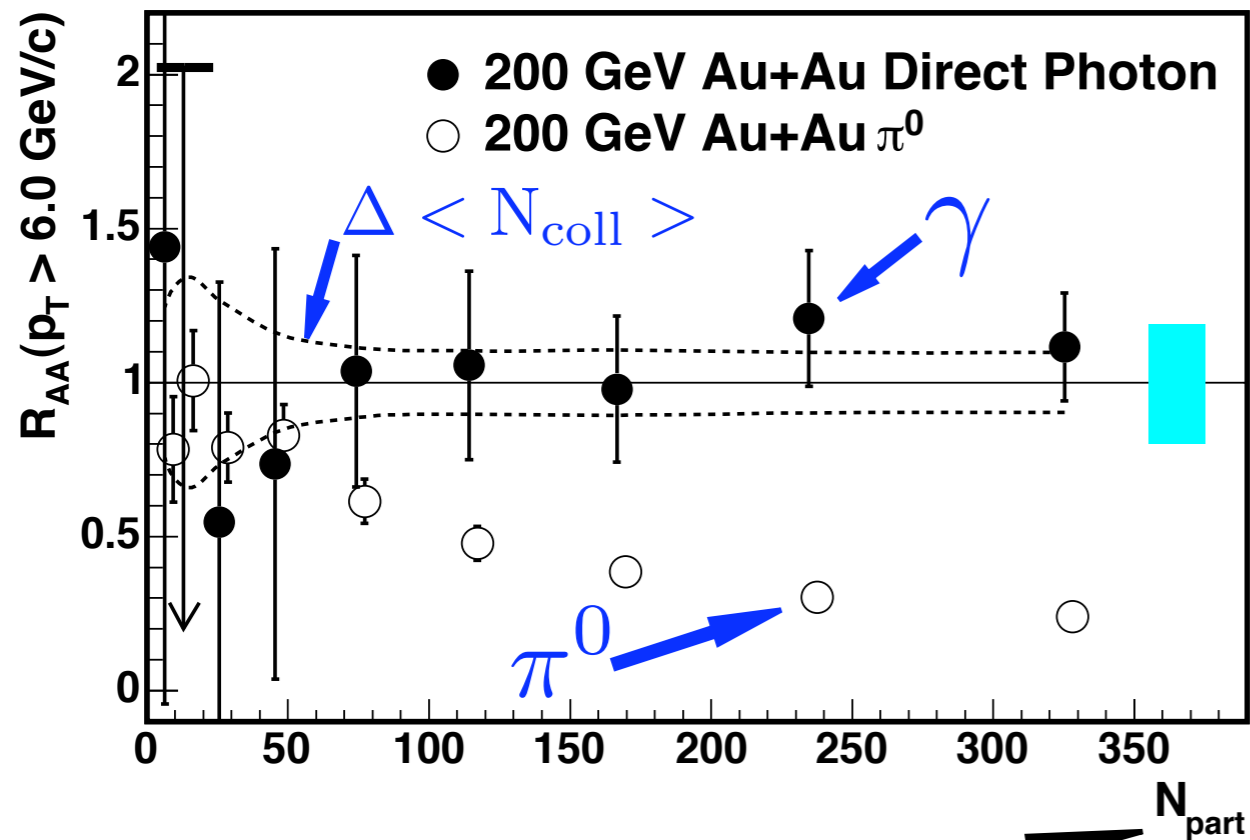
no sign of suppression of prompt  $\gamma$  in nuclear medium

# Comparison of direct photons with hadrons ( $\pi^0$ )

consider ratios of AA to pp yields scaled by  $N_{\text{coll}}$  for  $\gamma_{\text{dir}}$  and  $\pi^0$

$p_T > 6 \text{ GeV}$

$$R_{AA}(p_T) = \frac{(1/N_{AA}^{\text{evt}}) d^2 N_{AA}/dp_T dy}{\langle N_{\text{coll}} \rangle / \sigma_{pp}^{\text{inel}} \times d^2 \sigma_{pp}/dp_T dy}$$



participating nucleons



again from beam counters and zero deg calos

$\gamma$  not suppressed  
 $\pi^0$  suppression  
 increasing with  $N_{\text{part}}$

interpretation :

initial high  $p_T$   $\gamma$  creation unchanged

$\rightarrow \pi^0$  suppressed in final state

(hadron cloud transparent for gamma)

rules out explanations like gluon saturation in initial state

# Characteristic differences of discussed prompt $\gamma$ experiments

	reaction	hadronic energy[GeV]	d(vtx-calor) [m]	yield/bgd	isolation $R(\eta, \phi)$
HI ZEUS	$\gamma p, ep$	~200	1	shower analysis	1
D0	$p\bar{p}$	1800	1	shower analysis	0.4
CDF	$p\bar{p}$	1960	1.8	sh analysis +preshowr	0.4
CDF	$p\bar{p}$	1800	1.8	conversions	0.4
E607	$pp, pN$ $\pi N$	31, 39	9	measure $\gamma/\pi^0$	--
WA98	$Pb Pb$	17.3	22	measure $\gamma/\pi^0$	--
PHENIX	$Au Au$	200	5	measure $\gamma/\pi^0$	--

experiments exploiting  $\pi^0$  id, require no explicit isolation

impact on theory ? different contributions of fragmentation, soft gluons especially at high (RHIC) energies



# Questions concerning HERA results

Why NLO below the data (~30%)? Normalisation ok without hadronisation !

Is the hadronisation correction using PYTHIA/HERWIG too naive?

Can it reliably be applied to NLO partons, in particular for isolation cone?

Are there large ambiguities?

Are LO+parton shower models (PYTHIA, HERWIG) below data (~40%) for the same reason? Which?

They include hadronisation and should model cuts well.

Prompt  $\gamma$  + jet seem to be better described than inclusive prompt  $\gamma$ .

If so, is it that the jet requirements forces LO like configurations with less phase space for higher order effects?

# Questions to results in hadron-hadron reactions

Low  $\sqrt{s}$  ( $< 100$  GeV), pp or light nuclei :

Can good theoretical description at low  $p_t$  ( $< 5$  GeV) be achieved without ad hoc  $k_t$  smearing?

Can be stated, which data can be accommodated with theory and which not ?

Does  $k_t$  factorisation help (calculations a la Lipatov, Zotov) ?

High Energies :

Deviations from NLO seem to be smaller for preliminary D0 than CDF data.

Ways to clarification?

Heavy Nuclei :

How safe is the  $\sim 10\%$  uncertainty estimate of  $N_{\text{coll}}$  scaling ?

Can future RHIC data establish plasma effects at low  $p_t$  ?

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to summarize :

prompt photons provide  
an interesting pQCD laboratory,  
enlightening results, more to come